

## Non-Abelian geometric potentials and spin-orbit coupling for periodically driven systems

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We demonstrate the emergence of the non-Abelian geometric potentials and thus the three-dimensional (3D) spin-orbit coupling (SOC) for ultracold atoms without using the laser beams. This is achieved by subjecting an atom to a periodic perturbation which is the product of a position-dependent Hermitian operator  $\hat{V}(\mathbf{r})$  and a fast oscillating periodic function  $f(\omega t)$  with a zero average. To have a significant spin-orbit coupling, we analyze a situation where the characteristic energy of the periodic driving is not necessarily small compared to the driving energy  $\hbar\omega$ . Applying a unitary transformation to eliminate the original periodic perturbation, we arrive at a non-Abelian (noncommuting) vector potential term describing the 3D SOC. The general formalism is illustrated by analyzing the motion of an atom in a spatially inhomogeneous magnetic field oscillating in time. A cylindrically symmetric magnetic field provides the SOC involving the coupling between the spin  $\mathbf{F}$  and all three components of the orbital angular momentum (OAM)  $\mathbf{L}$ . In particular, the spherically symmetric monopole-type synthetic magnetic field  $\mathbf{B} \propto \mathbf{r}$  generates the 3D SOC of the  $\mathbf{L} \cdot \mathbf{F}$  form, which resembles the fine-structure interaction of hydrogen atom. However, the strength of the SOC here goes as  $1/r^2$  for larger distances, instead of  $1/r^3$  as in atomic fine structure. Such a longer-ranged SOC significantly affects not only the lower states of the trapped atom, but also the higher ones. Furthermore, by properly tailoring the external trapping potential, the ground state of the system can occur at finite OAM, while the ground state of hydrogen atom has zero OAM.

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### I. INTRODUCTION

The periodic driving enriches topological [1–11] and many-body [8,12–29] properties of physical systems. This can be used to generate the artificial gauge fields for ultracold atoms [30–42] and photonic systems [43–48], as well as to alter the topological properties of condensed-matter systems [2,4,49–55]. In many cases the periodic driving changes in time, which applies *inter alia* to experiments on ultracold atoms where the periodic driving is often slowly ramped up [27]. In such a situation the evolution of the system can be described in terms of a slowly changing effective Floquet Hamiltonian and a fast oscillating micromotion operator [56]. In particular, this is the case if the time periodic Hamiltonian is a product of a slowly changing operator  $\hat{V}(\lambda(t))$  and a fast oscillating function  $f(\omega t) = f(\omega t + 2\pi)$  with a zero average, where the vector  $\lambda(t)$  represents a set of slowly changing parameters [56,57]. If the operator  $\hat{V}(\lambda(t))$  does not commute with itself at different times, the effective evolution of the periodically driven system can be accompanied by non-Abelian (noncommuting) geometric phases after the vector  $\lambda(t)$  undergoes a cyclic change and returns to its original value [56–58].

Here we study a way to generate non-Abelian geometric potentials when the incident parameter  $\lambda(t)$  is replaced by a

radius vector of a particle  $\mathbf{r} = x\mathbf{e}_x + y\mathbf{e}_y + z\mathbf{e}_z$  representing a dynamical variable, and a kinetic-energy operator is added. If the operator  $\hat{V}(\mathbf{r})$  featured in the periodic coupling term  $\hat{V}(\mathbf{r})f(\omega t)$  does not commute with itself at different positions,  $[\hat{V}(\mathbf{r}), \hat{V}(\mathbf{r}')] \neq 0$ , the adiabatic evolution of the system within a Floquet band can be accompanied by a non-Abelian (noncommuting) geometric vector potential providing a three-dimensional (3D) spin-orbit coupling (SOC). The 2D and 3D SOC can be also generated optically by using degenerate eigenstates of the atom-light coupling operator known as dressed states [59–68]. This requires a considerable amount of effort [64,66–68]. Furthermore, the formation of the SOC is accompanied by unwanted heating due to the radiative decay of atoms in the dressed states.

The present approach does not rely on the degenerate atom-light dressed states. Instead, employment of the time-periodic interaction of the form  $\hat{V}(\mathbf{r})f(\omega t)$  provides degenerate Floquet states [57]. The spatial and temporal dependence of these states yields the non-Abelian vector potential and thus the 3D SOC. The general formalism is illustrated by analyzing the motion of an atom in a spatially inhomogeneous magnetic field oscillating in time. We study a cylindrically symmetric magnetic field and analyze the coupling between the atomic spin  $\mathbf{F}$  and the orbital angular momentum (OAM)  $\mathbf{L}$  for such a system. In particular, the monopole-type magnetic field generates the 3D SOC of the  $\mathbf{L} \cdot \mathbf{F}$  form involving the coupling between the atomic spin  $\mathbf{F}$  and OAM.

We show that the strength of the  $\mathbf{L} \cdot \mathbf{F}$  SOC is long ranged and goes as  $1/r^2$  for larger distances, rather than as  $1/r^3$ , as experienced by an electron in a Coulomb potential via the fine-structure interaction [69]. For such larger distances

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exceeding a characteristic SOC range  $r_0$ , the SOC contribution reduces to  $-\mathbf{L}^2/2mr^2$  and thus it cancels the centrifugal term featured in the kinetic-energy operator. Therefore, the SOC significantly affects all atomic states. In the case of the (quasi)spin 1/2 atom, the SOC makes the atomic states nearly degenerate with respect to the orbital quantum number  $l$  for a fixed total angular momentum quantum number  $j = l \pm 1/2$  and a fixed radial quantum number  $n_r$ . Furthermore, for a harmonic trap the ground state with  $j = 1/2$  and  $l = 0$  has a slightly lower energy than the one with  $j = 1/2$  and  $l = 1$ . The situation can be changed by adding an extra antitrapping potential for small  $r$ . In that case the ground state of the system acquires a nonzero orbital quantum number  $l = 1$  and thus is affected by the SOC. This is a consequence of the periodic driving; normally the ground state of a spherically symmetric SOC system corresponds to  $l = 0$ .

The paper is organized as follows. The general formalism is presented in the subsequent Sec. II. We define a periodically driven system and apply a unitary transformation eliminating the time-oscillating part of the original Hamiltonian. The evolution of the transformed state vector is then governed by a new Hamiltonian  $W(\omega t)$  containing a vector potential type contribution  $\hat{\mathbf{A}}(\mathbf{r}, \omega t)$  which gives the spin-dependent momentum shift and provides the SOC in the effective Floquet Hamiltonian. We also present a general analysis of the vector potential and discuss ramping of the periodic driving. In Sec. III, as a specific example, we study the spin in an oscillating magnetic field. We present an explicit expression for the time-periodic vector potential, and consider the coupling between the spin and OAM for a cylindrically symmetric magnetic field. In Sec. IV we analyze the SOC for the spherically symmetric monopole magnetic field. Section V considers the adiabatic condition and discusses possible experimental implementations. Section VI presents concluding remarks. Details of some technical calculations are contained in two Appendixes A and B. In particular, in Appendix A we present a way of producing the interaction of an atom with an effective magnetic field violating the Maxwell equations, including the monopole magnetic field [70].

## II. PERIODICALLY DRIVEN SYSTEM WITH A MODULATED DRIVING

### A. Hamiltonian and equations of motion

Let us consider the center-of-mass motion of a quantum particle such as an ultracold atom. The particle is subjected to a periodic driving described by the operator  $\hat{V}(\mathbf{r})f(\omega t)$ , which is a product of a position-dependent Hermitian operator  $\hat{V}(\mathbf{r})$  and a time-periodic function  $f(\omega t + 2\pi) = f(\omega t)$  with a zero average  $\int_{-\pi}^{\pi} f(\omega t) dt = 0$ . A hat over the operator  $\hat{V}(\mathbf{r})$  indicates that it depends on the internal degrees of freedom of the particle. The state-dependent operator  $\hat{V}(\mathbf{r})$  generally does not commute with itself at different positions  $[\hat{V}(\mathbf{r}), \hat{V}(\mathbf{r}')] \neq 0$ . Including the kinetic energy, the system is described by the time-periodic Hamiltonian  $\hat{H}(\omega t) = \hat{H}(\omega t + 2\pi)$  given by

$$\hat{H}(\omega t) = \frac{\mathbf{p}^2}{2m} + \hat{V}(\mathbf{r})f(\omega t) + V_{\text{ex}}(\mathbf{r}), \quad (1)$$

where  $\mathbf{p} = -i\hbar\nabla$  is the momentum operator,  $m$  is the mass of the particle, and we have also added an extra

potential  $V_{\text{ex}}(\mathbf{r})$  to confine the particle in a trap. The external potential is considered to be state independent, so that  $[V_{\text{ex}}(\mathbf{r}), V_{\text{ex}}(\mathbf{r}')] = 0$ .

The system is described by a state vector  $|\phi(t)\rangle$  obeying the time-dependent Schrödinger equation (TDSE):

$$i\hbar \frac{\partial}{\partial t} |\phi(t)\rangle = \hat{H}(\omega t) |\phi(t)\rangle. \quad (2)$$

An example of such a system is an atom in a spatially inhomogeneous magnetic field  $\mathbf{B}(\mathbf{r})f(\omega t)$  with a fast oscillating amplitude  $\propto f(\omega t)$  and a slowly changing magnitude or direction of the amplitude  $\mathbf{B}(\mathbf{r})$ . In that case the position-dependent part of the operator  $\hat{V}(\mathbf{r})f(\omega t)$  is given by

$$\hat{V}(\mathbf{r}) = g_F \hat{\mathbf{F}} \cdot \mathbf{B}(\mathbf{r}), \quad (3)$$

where  $g_F$  is a gyromagnetic factor and  $\hat{\mathbf{F}} = \hat{F}_1 \mathbf{e}_x + \hat{F}_2 \mathbf{e}_y + \hat{F}_3 \mathbf{e}_z$  is the spin operator with the Cartesian components obeying the commutation relations  $[\hat{F}_s, \hat{F}_q] = i\hbar \epsilon_{sqr} \hat{F}_r$ . Here  $\epsilon_{sqr}$  is a Levi-Civita symbol and the summation over a repeated Cartesian index  $u = x, y, z$  is implied. The operator  $\hat{V}(\mathbf{r})$  does not commute with itself at different positions  $[\hat{V}(\mathbf{r}), \hat{V}(\mathbf{r}')] \neq 0$  if  $\mathbf{B}(\mathbf{r})$  and  $\mathbf{B}(\mathbf{r}')$  are oriented along different axes. This leads to the SOC for the spin in the spatially nonuniform magnetic field oscillating in time, to be studied in the subsequent Secs. III–V.

### B. Transformed representation

To have a significant SOC, we consider a situation where the matrix elements of the periodic driving  $\hat{V}(\mathbf{r})f(\omega t)$  are not necessarily small compared to the driving energy  $\hbar\omega$ . In that case one cannot apply the high frequency expansion [56,71,72] of an effective Floquet Hamiltonian in the original representation. To bypass the problem, we go to a new representation via a unitary transformation eliminating the operator  $\hat{V}(\mathbf{r})f(\omega t)$ :

$$\hat{R} = \hat{R}(\mathbf{r}, \omega t) = \exp \left[ -i \frac{\mathcal{F}(\omega t)}{\hbar\omega} \hat{V}(\mathbf{r}) \right], \quad (4)$$

where  $\mathcal{F}(\theta)$  is a primitive function of  $f(\theta) = d\mathcal{F}(\theta)/d\theta$  with a zero average  $[\int_{-\pi}^{\pi} \mathcal{F}(\theta') d\theta' = 0]$  and a calligraphy letter  $\mathcal{F}$  is used to avoid confusion with the spin operator  $\mathbf{F}$  featured in Eq. (3). Since  $\mathcal{F}(\omega t) = \mathcal{F}(\omega t + 2\pi)$ , the transformation  $\hat{R}(\mathbf{r}, \omega t) = \hat{R}(\mathbf{r}, \omega t + 2\pi)$  has the same periodicity as the original Hamiltonian  $\hat{H}(\omega t) = \hat{H}(\omega t + 2\pi)$ .

The transformed state vector

$$|\psi(t)\rangle = \hat{R}^\dagger(\mathbf{r}, \omega t) |\phi(t)\rangle \quad (5)$$

obeys the TDSE

$$i\hbar \frac{\partial}{\partial t} |\psi(t)\rangle = \hat{W}(\omega t) |\psi(t)\rangle \quad (6)$$

governed by the Hamiltonian

$$\hat{W}(\omega t) = \frac{1}{2m} [\mathbf{p} - \hat{\mathbf{A}}(\mathbf{r}, \omega t)]^2 + V_{\text{ex}}(\mathbf{r}), \quad (7)$$

where a time-periodic vector potential type operator

$$\hat{\mathbf{A}}(\mathbf{r}, \omega t) = i\hbar \hat{R}^\dagger(\mathbf{r}, \omega t) \nabla \hat{R}(\mathbf{r}, \omega t) \quad (8)$$

is added to the momentum operator  $\mathbf{p}$  due to the position dependence of the unitary transformation  $\hat{R}(\mathbf{r}, \omega t)$ . On the

other hand, the transformation  $\hat{R}(\mathbf{r}, \omega t)$  does not affect the state-independent trapping potential  $V_{\text{ex}}(\mathbf{r})$ . The transformed Hamiltonian  $\hat{W}(\omega t)$  no longer contains the time-periodic term  $\hat{V}(\mathbf{r})f(\omega t)$ . The periodic driving is now represented by the operator  $\hat{\mathbf{A}}(\mathbf{r}, \omega t) = \hat{\mathbf{A}}(\mathbf{r}, \omega t + 2\pi)$  featured in the transformed Hamiltonian (7). This leads to the SOC to be studied in the next subsection.

In this way the properly chosen transformation  $\hat{R}(\mathbf{r}, \omega t)$  eliminates the interaction operator  $V(r)f(\omega t)$  in the original Hamiltonian (1). The position dependence of  $\hat{R}(\mathbf{r}, \omega t)$  yields the spin-dependent momentum shift  $\hat{\mathbf{A}}(\mathbf{r}, \omega t)$  in Eq. (7), so the SOC appears directly from the unitary transformation.

### C. Floquet adiabatic approach

The transformed Hamiltonian  $\hat{W}(\omega t) = \hat{W}(\omega t + 2\pi)$  can be expanded in the Fourier components:

$$\hat{W}(\omega t) = \sum_{n=-\infty}^{\infty} \hat{W}^{(n)} e^{in\omega t}, \quad (9)$$

with

$$\hat{W}^{(n)} = \frac{1}{2\pi} \int_{-\pi}^{\pi} \hat{W}(\theta) e^{-in\theta} d\theta. \quad (10)$$

In what follows the driving energy  $\hbar\omega$  is assumed to be much larger than the matrix elements of the Fourier components of the transformed Hamiltonian  $\hat{W}^{(n)}$ ,

$$\hbar\omega \gg |\hat{W}_{\alpha\beta}^{(n)}|, \quad (11)$$

where the superscript ( $n$ ) refers to the  $n$ th Fourier component. The condition (11) allows one to consider the adiabatic evolution of the system in a selected Floquet band by neglecting the nonzerth (with  $n \neq 0$ ) Fourier components  $\hat{W}^{(n)}$  of the transformed Hamiltonian  $\hat{W}(\omega t)$ . Thus one replaces the exact evolution governed by the time-dependent transformed Hamiltonian  $\hat{W}(\omega t)$  by the approximate one governed by the time-independent effective Floquet Hamiltonian  $\hat{W}_{\text{eff}(0)} = \hat{W}^{(0)}$  equal to the zeroth Fourier component:

$$i\hbar \frac{\partial}{\partial t} |\psi^{(0)}(t)\rangle = \hat{W}^{(0)} |\psi^{(0)}(t)\rangle, \quad (12)$$

where  $|\psi^{(0)}(t)\rangle$  is the corresponding approximate state vector representing the slowly changing part of the exact state vector  $|\psi(t)\rangle$ . The slowly changing state vector  $|\psi^{(0)}(t)\rangle$  deviates little from the exact time evolution of the state vector  $|\psi(t)\rangle$  if the adiabatic condition (11) holds.

The effective Floquet Hamiltonian corresponding to the transformed Hamiltonian (7) reads

$$\hat{W}^{(0)} = \frac{\mathbf{p}^2}{2m} + \hat{W}_{\text{SOC}} + \hat{V}_{\text{total}}(\mathbf{r}), \quad (13)$$

where

$$\hat{V}_{\text{total}}(\mathbf{r}) = \frac{1}{2m} \langle [\hat{\mathbf{A}}(\mathbf{r}, \omega t)]^2 \rangle + V_{\text{ex}}(\mathbf{r}) \quad (14)$$

is the total scalar potential and

$$\hat{W}_{\text{SOC}} = -\frac{1}{2m} [\hat{\mathbf{A}}^{(0)}(\mathbf{r}) \cdot \mathbf{p} + \mathbf{p} \cdot \hat{\mathbf{A}}^{(0)}(\mathbf{r})] \quad (15)$$

describes the SOC emerging via the zeroth Fourier component of the oscillating vector potential:  $\hat{\mathbf{A}}^{(0)}(\mathbf{r}) = \langle \hat{\mathbf{A}}(\mathbf{r}, \omega t) \rangle =$

$\frac{1}{2\pi} \int_{-\pi}^{\pi} \hat{\mathbf{A}}(\mathbf{r}, \theta) d\theta$ . Here the brackets  $\langle \dots \rangle$  signify the zero-frequency component (the time average) of an oscillating operator. In this way, the effective Hamiltonian  $\hat{W}^{(0)}$  is determined by the time averages of the oscillating vector potential  $\hat{\mathbf{A}}(\mathbf{r}, \omega t)$  and its square. The vector potential  $\hat{\mathbf{A}}^{(0)}(\mathbf{r})$  generally contains three noncommuting Cartesian components leading to the 3D SOC. Note that the Floquet adiabatic approach applied here corresponds to the zero order of the high frequency expansion [57,71–73] of the effective Floquet Hamiltonian  $W_{\text{eff}} = \hat{W}^{(0)} + O(1/\omega)$ .

The present perturbation analysis relies on the condition (11) involving the Fourier components of the time-periodic operator  $\hat{\mathbf{A}}(\mathbf{r}, \omega t) = \hat{\mathbf{A}}(\mathbf{r}, \omega t + 2\pi)$  given by Eq. (8). The operator  $\hat{\mathbf{A}}(\mathbf{r}, \omega t)$  emerges via the  $\mathbf{r}$  dependence of the ratio  $\frac{\mathcal{F}(\omega t)}{\hbar\omega} \hat{V}(\mathbf{r})$  featured in the exponent of the unitary transformation  $\hat{R}(\mathbf{r}, \omega t)$ . Therefore, the operator  $\hat{\mathbf{A}}(\mathbf{r}, \omega t)$  is determined by the spatial changes of the ratio  $\hat{V}(\mathbf{r})/\omega$ , and the Floquet adiabatic condition (11) requires the smallness of the spatial changes of the operator  $\hat{V}(\mathbf{r})$  rather than the smallness of the operator  $\hat{V}(\mathbf{r})$  itself with respect to the driving frequency  $\omega$ .

The condition (11) can hold even if the matrix elements of the periodic driving  $\hat{V}(\mathbf{r})f(\omega t)$  are not small compared to the driving energy  $\hbar\omega$ . In such a situation the high frequency expansion of the effective Hamiltonian [57,71–73] is not applicable in the original representation where the evolution is given by Eq. (2). Yet it is applicable in the transformed representation corresponding to the equation of motion (6). Therefore, the present approach allows one to realize the SOC which is much larger than the one relying on the perturbation treatment in the original representation, as it was done in a very recent study [74]. The adiabatic condition (11) will be analyzed in more detail for a spin in an oscillating magnetic field in Sec. V A.

Returning to the original representation, the adiabatic evolution of the state vector is given by

$$|\phi(t)\rangle \equiv |\phi(\omega t, t)\rangle \approx \hat{R}(\mathbf{r}, \omega t) |\psi^{(0)}(t)\rangle. \quad (16)$$

Since  $\hat{R}(\mathbf{r}, \omega t) = \hat{R}(\mathbf{r}, \omega t + 2\pi)$ , the original state vector  $|\phi(\omega t, t)\rangle = |\phi(\omega t + 2\pi, t)\rangle$  is  $2\pi$  periodic with respect to the first variable. Therefore,  $\hat{R}(\mathbf{r}, \omega t)$  describes the fast micromotion of the original state vector  $|\phi(\omega t, t)\rangle$ . Additionally, the state vector  $|\phi(\omega t, t)\rangle$  changes slowly with respect to the second variable due to the slow changes of the transformed state vector  $|\psi^{(0)}(t)\rangle$ .

### D. Equation for $\hat{\mathbf{A}}(\mathbf{r}, \omega t)$ and its expansion

To obtain an equation for the vector potential  $\hat{\mathbf{A}} = \hat{\mathbf{A}}(\mathbf{r}, \omega t)$ , let us treat it as a function of the coordinate  $\mathbf{r}$  and a parameter  $c = c(\omega t) \equiv \mathcal{F}(\omega t)/\hbar\omega$ . Differentiating  $\hat{\mathbf{A}}(\mathbf{r}, \omega t) = \hat{\mathbf{A}}(\mathbf{r}; c)$  with respect to  $c$  for fixed  $\mathbf{r}$ , and using Eqs. (4) and (8), one arrives at the following differential equation for the Cartesian components  $\hat{A}_u$  of the vector potential:

$$\frac{\partial \hat{A}_u}{\partial c} = \hbar \frac{\partial \hat{V}}{\partial u} + i[\hat{V}, \hat{A}_u] \quad (u = x, y, z), \quad (17)$$

subject to the initial condition

$$\hat{A}_u = 0 \quad \text{for } c = 0. \quad (18)$$

A solution to Eq. (17) can be expanded in the powers of  $c \propto 1/\omega$ , giving

$$\begin{aligned} \hat{\mathbf{A}}(\mathbf{r}, \omega t) &= \frac{\mathcal{F}(\omega t)}{\omega} \nabla \hat{V} + i \frac{\mathcal{F}^2(\omega t)}{2! \hbar \omega^2} [\hat{V}, \nabla \hat{V}] \\ &+ i^2 \frac{\mathcal{F}^3(\omega t)}{3! \hbar^2 \omega^3} [\hat{V}, [\hat{V}, \nabla \hat{V}]] + \dots \end{aligned} \quad (19)$$

If for any  $\mathbf{r}$  and  $\mathbf{r}'$  the commutator  $[\hat{V}(\mathbf{r}), \hat{V}(\mathbf{r}')] = 0$ , then only the first term remains in the expansion (19):

$$\hat{\mathbf{A}}(\mathbf{r}, \omega t) = \frac{\mathcal{F}(\omega t)}{\omega} \nabla \hat{V}, \quad (20)$$

giving  $\langle \hat{\mathbf{A}}(\mathbf{r}, \omega t) \rangle = \hat{\mathbf{A}}^{(0)}(\mathbf{r}) = 0$ . In that case no SOC is generated, and the time average  $\langle [\hat{\mathbf{A}}(\mathbf{r}, \omega t)]^2 \rangle$  provides an extra trapping potential in Eq. (14). In particular, this applies to a state-independent potential  $\hat{V}(\mathbf{r}) = V(\mathbf{r})$  for which the time-periodic Hamiltonian (1) describes the Kapitza problem [73].

Here we go beyond the situation where  $[\hat{V}(\mathbf{r}), \hat{V}(\mathbf{r}')] = 0$ , so the commutators are nonzero in the expansion (19). As a result, the vector potential  $\hat{\mathbf{A}}(\mathbf{r}, \omega t)$  has a nonzero average  $\hat{\mathbf{A}}^{(0)}(\mathbf{r}) \neq 0$  providing the SOC which acts in all three dimensions. Such a 3D SOC can be realized for a spinful atom in an inhomogeneous magnetic field oscillating in time, to be considered in Sec. III. In that case the vector potential  $\hat{\mathbf{A}}(\mathbf{r}, \omega t)$  is obtained exactly in Eq. (24) which is valid for an arbitrary driving frequency, not necessarily small compared to the strength of the periodic driving. Thus the solution (24) effectively takes into account all the terms in the expansion of  $\hat{\mathbf{A}}(\mathbf{r}, \omega t)$  given by Eq. (19).

### E. Ramping of the periodic perturbation

Up to now the operator  $\hat{V}(\mathbf{r})$  defining the periodic driving in the Hamiltonian (1) was considered to be time independent, so the driving was strictly periodic in time. The analysis can be extended to a situation where the operator  $\hat{V}(\mathbf{r})$  has an extra slow temporal dependence [56–58]. This can describe ramping of the periodic perturbation. It is quite common to have no periodic driving at an initial time  $t = t_{\text{in}}$  and ramp up the driving slowly afterwards over a time much larger than the driving period  $T = 2\pi/\omega$ . This can be described by a slowly changing factor  $\alpha(t)$  multiplying  $\hat{V}(\mathbf{r})$ :

$$\hat{V}(\mathbf{r}, \alpha(t)) = \alpha(t) \hat{V}(\mathbf{r}), \quad (21)$$

where  $\alpha(t)$  changes smoothly from  $\alpha(t) = 0$  at the initial time  $t = t_{\text{in}}$  to  $\alpha(t) = 1$  at the final stage of the ramping. In particular, Eq. (21) with  $\hat{V}(\mathbf{r})$  given by Eq. (3) describes a spin in an oscillating magnetic field with a slowly ramped amplitude  $\mathbf{B}(\mathbf{r}, \alpha(t)) = \alpha(t) \mathbf{B}(\mathbf{r})$ .

The slow temporal dependence of  $\hat{V}(\mathbf{r}, \alpha(t))$  featured in the unitary transformation  $\hat{R}$  provides an additional term  $\hat{W}_{\text{add}}(\mathbf{r}, \omega t, t)$  to the transformed Hamiltonian  $\hat{W}(\omega t, t)$  [57]:

$$\hat{W}_{\text{add}}(\mathbf{r}, \omega t, t) = -i \hbar \dot{\alpha} \hat{R}^\dagger(\mathbf{r}, \omega t, \alpha) \frac{\partial \hat{R}(\mathbf{r}, \omega t, \alpha)}{\partial \alpha}. \quad (22)$$

The operator  $\hat{V}(\mathbf{r}, \alpha(t))$  commutes with itself at different times,  $[\hat{V}(\mathbf{r}, \alpha(t)), \hat{V}(\mathbf{r}, \alpha(t'))] = 0$ , giving

$$\hat{W}_{\text{add}}(\mathbf{r}, \omega t, t) = -\frac{\mathcal{F}(\omega t) \dot{\alpha}}{\omega} \frac{\partial \hat{V}(\mathbf{r}, \alpha)}{\partial \alpha}. \quad (23)$$

Since  $\langle \mathcal{F}(\omega t) \rangle = 0$ , the extra term  $\hat{W}_{\text{add}}(\mathbf{r}, \omega t, t)$  averages to zero and thus has no zero Fourier component  $\hat{W}_{\text{add}}^{(0)}(\mathbf{r}, \omega t, t) = 0$ . In this way, the ramping of the periodic driving described by Eq. (21) does not provide an extra contribution to the effective Hamiltonian and thus does not affect the effective dynamics of the system.

## III. SPIN IN TIME-OSCILLATING MAGNETIC FIELD

### A. Vector potential

The general formalism is illustrated by considering motion of a spinful atom in a time-oscillating magnetic field with the interaction operator  $\hat{V}(\mathbf{r})$  given by Eq. (3). In that case the operator  $\hat{\mathbf{A}} = \hat{\mathbf{A}}(\mathbf{r}, \omega t)$  can be derived exactly for an arbitrary strength of the magnetic field. Specifically, by solving Eq. (17) one arrives at the following Cartesian components of the vector potential  $\hat{\mathbf{A}}(\mathbf{r}, \omega t)$ :

$$\begin{aligned} \hat{A}_u(\mathbf{r}, \omega t) &= a \mathcal{F} \frac{(\mathbf{B} \cdot \partial \mathbf{B} / \partial u)(\mathbf{B} \cdot \hat{\mathbf{F}})}{B^3} \\ &+ \sin(a \mathcal{F}) \frac{[(\mathbf{B} \times \partial \mathbf{B} / \partial u) \times \mathbf{B}] \cdot \hat{\mathbf{F}}}{B^3} \\ &+ [\cos(a \mathcal{F}) - 1] \frac{(\mathbf{B} \times \partial \mathbf{B} / \partial u) \cdot \hat{\mathbf{F}}}{B^2}, \end{aligned} \quad (24)$$

where

$$a = \frac{B g_F}{\omega} \quad (25)$$

defines the frequency of the magnetic interaction in the units of the driving frequency. Here we keep implicit the time dependence of the oscillating function  $\mathcal{F} = \mathcal{F}(\omega t)$ , as well as the  $\mathbf{r}$  dependence of  $\mathbf{B} = \mathbf{B}(\mathbf{r})$  and  $a = a(\mathbf{r})$ . The derivation of Eq. (24) is analogous to the one presented in the Appendix of Ref. [57] subject to replacement of the time derivatives  $\partial V / \partial t$  and  $\partial \mathbf{B} / \partial t$  by the space derivatives  $-\partial V / \partial u$  and  $-\partial \mathbf{B} / \partial u$ , respectively.

### B. Time-averaged vector potential and SOC term

To simplify the subsequent analysis, we assume the original Hamiltonian (1) to have a time-reversal symmetry. This is the case if the function  $f(\omega t)$  describing the periodic driving is even:  $f(\omega t) = f(-\omega t)$  (subject to a proper choice of the origin of time). Consequently the function  $\mathcal{F}(\omega t) = \omega \int_0^t f(\omega t') dt'$  featured in the vector potential  $\hat{\mathbf{A}}(\mathbf{r}, \omega t)$  is an odd function:  $\mathcal{F}(\omega t) = -\mathcal{F}(-\omega t)$ . In particular, this holds for a harmonic driving with

$$f(\omega t) = \cos(\omega t), \quad \mathcal{F}(\omega t) = \sin(\omega t). \quad (26)$$

For  $\mathcal{F}(\omega t) = -\mathcal{F}(-\omega t)$  the first two lines of Eq. (24) are odd functions of time and thus average to zero, giving

$$\hat{A}_u^{(0)}(\mathbf{r}) = \frac{\langle \cos(a \mathcal{F}) \rangle - 1}{B^2} (\mathbf{B} \times \partial \mathbf{B} / \partial u) \cdot \hat{\mathbf{F}}. \quad (27)$$

Note that, for the harmonic driving (26), the time average  $\langle \cos(a \mathcal{F}) \rangle$  is given by the Bessel functions of the first kind:

$$\langle \cos(a \mathcal{F}) \rangle = \mathcal{J}_0(a), \quad \text{with} \quad \mathcal{J}_0(a) = \frac{1}{2\pi} \int_{-\pi}^{\pi} e^{ia \sin \theta} d\theta. \quad (28)$$

Substituting Eq. (27) into (15), the SOC term takes the form

$$\hat{W}_{\text{SOC}} = \frac{1}{2m} \left[ \frac{1 - \langle \cos(a\mathcal{F}) \rangle}{B^2} \mathbf{Z} \cdot \hat{\mathbf{F}} + \mathbf{Z}^\dagger \cdot \hat{\mathbf{F}} \frac{1 - \langle \cos(a\mathcal{F}) \rangle}{B^2} \right], \quad (29)$$

where

$$\mathbf{Z} = \mathbf{B} \times \sum_{u=x,y,z} \frac{\partial \mathbf{B}}{\partial u} p_u \quad (30)$$

is an orbital operator, which makes it clear that Eq. (29) represents SOC.

In addition to the SOC term  $\hat{W}_{\text{SOC}}$ , the effective Hamiltonian (13) contains also the  $\langle [\hat{\mathbf{A}}(\mathbf{r}, \omega t)]^2 \rangle$  term which is featured in scalar potential  $V_{\text{total}}(r)$  given by Eq. (14). This contribution is analyzed in Appendix B for specific configurations of the magnetic field.

The form of the SOC term in Eq. (29) is valid for an arbitrary magnetic configuration. However, to give a concrete example and to make the physics more clear, in the next subsection we consider a cylindrically symmetric magnetic field given by Eq. (31). Subsequently in Sec. IV we analyze an important particular situation where the magnetic field takes the form of a monopole, in which case we can take advantage of the spherical symmetry of the system and simplify the calculations. In Appendix A we present details on how to generate various effective magnetic fields with a nonzero divergence field including the cylindrically symmetric magnetic field (31) and the effective monopole field (33).

### C. Cylindrical magnetic field

#### 1. Magnetic field

Let us consider the magnetic field  $\mathbf{B} = \mathbf{B}(\mathbf{r})$ , which changes linearly in space and has a cylindric symmetry:

$$\mathbf{B}(\mathbf{r}) = \alpha_\perp(x\mathbf{e}_x + y\mathbf{e}_y) + \alpha_z z\mathbf{e}_z, \quad (31)$$

where the ratio between  $\alpha_z$  and  $\alpha_\perp$  is considered to be arbitrary. By taking  $\alpha_z = -2\alpha_\perp$ , Eq. (31) describes a quadrupole magnetic field [75]. On the other hand, for  $\alpha_z \neq -2\alpha_\perp$  the cylindrically symmetric magnetic field has a nonzero divergence and thus does not obey the Maxwell equation  $\nabla \cdot \mathbf{B}(\mathbf{r}) = 0$ . In particular, this is the case for a spherically symmetric monopole field corresponding to  $\alpha_z = \alpha_\perp = \alpha$  and considered in Sec. IV.

#### 2. SOC operator

The operator  $\mathbf{Z} \cdot \hat{\mathbf{F}}$  entering Eq. (29) for the SOC operator  $\hat{W}_{\text{SOC}}$  reads for the cylindrically symmetric magnetic field (31)

$$\mathbf{Z} \cdot \hat{\mathbf{F}} = \mathbf{Z}^\dagger \cdot \hat{\mathbf{F}} = \alpha_\perp^2 L_z \hat{F}_z + \alpha_\perp \alpha_z (L_x \hat{F}_x + L_y \hat{F}_y), \quad (32)$$

where  $L_x$ ,  $L_y$ , and  $L_z$  are the Cartesian components of the OAM operator  $\mathbf{L} = \mathbf{r} \times \mathbf{p}$ . Note that the function  $[1 - \langle \cos(a\mathcal{F}) \rangle]/B^2$  featured in Eq. (29) is cylindrically symmetric and thus preserves the  $z$  projection of the OAM for the magnetic field (31).

The term  $L_z \hat{F}_z$  in Eq. (32) provides the spin-dependent shift to eigenenergies of the OAM operator  $L_z$ . On the other hand, the term  $L_x \hat{F}_x + L_y \hat{F}_y = L_+ \hat{F}_- + L_- \hat{F}_+$  represents transitions

between different spin and OAM projection states described by the raising or lowering operators  $L_\pm = L_x \pm iL_y$  and  $\hat{F}_\pm = \hat{F}_x \pm i\hat{F}_y$ . Therefore, the present SOC has some similarities to the coupling between the spin and OAM induced by Raman laser beams carrying optical vortices [76–87]. Yet, unlike the Raman case, now the coupling between the spin and OAM is described by all three OAM projections  $L_x$ ,  $L_y$ , and  $L_z$  as long as  $\alpha_\perp \neq 0$  and  $\alpha_z \neq 0$ , so the coupling is truly three dimensional. In particular, for a monopole magnetic field where  $\alpha_\perp = \alpha_z = \alpha$ , Eqs. (29) and (32) yield a spherically symmetric coupling between the spin and OAM  $\propto \mathbf{L} \cdot \hat{\mathbf{F}}$  presented by Eq. (34) below.

## IV. MONOPOLE FIELD

### A. Effective Hamiltonian for monopole field

For  $\alpha_z = \alpha_\perp = \alpha$ , Eq. (31) reduces to the centrally symmetric monopolelike magnetic field

$$\mathbf{B} = \alpha \mathbf{r} = \frac{2\omega}{r_0 g_F} \mathbf{r}, \quad (33)$$

where  $r_0 = 2\omega/\alpha g_F$  defines the radius  $r = r_0$  at which a characteristic frequency of the magnetic interaction  $g_F B/2 = \omega r/r_0$  becomes equal to the driving frequency  $\omega$ . In such a situation, the operator  $\mathbf{Z} \cdot \hat{\mathbf{F}} = \alpha^2 \mathbf{L} \cdot \hat{\mathbf{F}}$  commutes with the spherically symmetric magnetic field  $B = \alpha r$ , so ordering of operators is not important in the SOC term (29), giving

$$\hat{W}_{\text{SOC}} = \hbar \omega_{\text{SOC}}(r) \frac{\mathbf{L} \cdot \hat{\mathbf{F}}}{\hbar^2}, \quad (34)$$

where the frequency

$$\omega_{\text{SOC}}(r) = \frac{\hbar}{m} \frac{1 - \langle \cos(2r\mathcal{F}/r_0) \rangle}{r^2} \quad (35)$$

characterizes the SOC strength. On the other hand, the term  $\langle [\hat{\mathbf{A}}(\mathbf{r}, \omega t)]^2 \rangle$  featured in the total scalar potential  $V_{\text{total}}(r)$  is given by Eq. (B8) in Appendix B. Combining Eqs. (13), (14), (34), (35), and (B8), the effective Hamiltonian takes the form

$$\hat{W}^{(0)} = \frac{p^2}{2m} + \hbar \omega_{\text{SOC}}(r) \left[ \frac{\mathbf{L} \cdot \hat{\mathbf{F}}}{\hbar^2} + \frac{r^2 \mathbf{F}^2 - (\mathbf{r} \cdot \hat{\mathbf{F}})^2}{\hbar^2 r^2} \right] + \frac{2}{mr_0^2} \langle \mathcal{F}^2 \rangle \frac{(\mathbf{r} \cdot \hat{\mathbf{F}})^2}{r^2} + V_{\text{ex}}(\mathbf{r}), \quad (36)$$

where the spin-dependent operator  $(\mathbf{r} \cdot \hat{\mathbf{F}})^2$  emerges from the  $\langle [\hat{\mathbf{A}}(\mathbf{r}, \omega t)]^2 \rangle$  term entering the total scalar potential (14). Generally the operator  $(\mathbf{r} \cdot \hat{\mathbf{F}})^2$  does not commute with  $\mathbf{L}^2$ , and thus mixes the states with different orbital quantum numbers  $l$ . Specifically, the term  $(\mathbf{r} \cdot \hat{\mathbf{F}})^2$  can provide coupling between orbit and spin or even spin tensor involving the radius vector  $\mathbf{r}$  rather than the momentum operator, as in Ref. [88]. However, no such extra SOC appears for the spin-1/2 atom to be considered next.

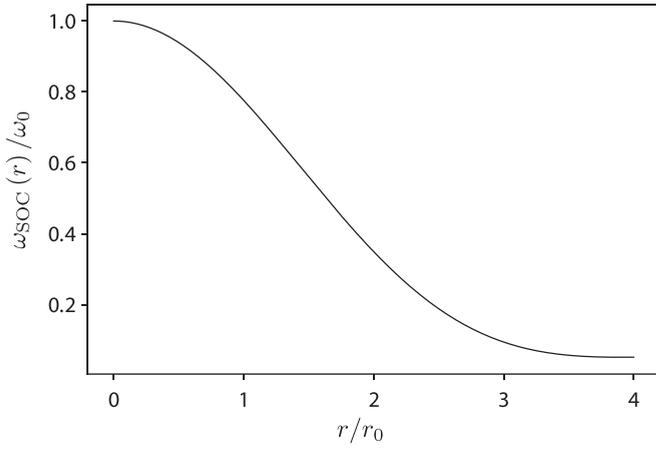


FIG. 1. Radial dependence of the SOC energy  $\omega_{\text{SOC}}(r)$  given by Eq. (35) for the sinusoidal driving (26) for which  $\langle \cos(2r\mathcal{F}/r_0) \rangle = \mathcal{J}_0(2r/r_0)$ . The distance is measured in the units of  $r_0$  and the frequency is measured in the units of the SOC frequency  $\omega_0 = \omega_{\text{SOC}}(0)$  given by Eq. (44).

## B. Spin 1/2

### 1. Effective Hamiltonian

Let us now consider the effective Hamiltonian (36) for a spin-1/2 atom for which

$$\hat{\mathbf{F}} = \frac{\hbar}{2} \hat{\boldsymbol{\sigma}}, \quad (37)$$

where  $\hat{\boldsymbol{\sigma}} = \hat{\sigma}_x \mathbf{e}_x + \hat{\sigma}_y \mathbf{e}_y + \hat{\sigma}_z \mathbf{e}_z$  and  $\hat{\sigma}_{x,y,z}$  are the Pauli matrices. In that case the operator  $(\mathbf{r} \cdot \hat{\mathbf{F}})^2 = \hbar^2 r^2 / 4$  is spin independent and spherically symmetric, making  $\mathbf{L}^2$  a conserving quantity. Using  $2\mathbf{L} \cdot \hat{\mathbf{F}} = \hat{\mathbf{J}}^2 - \mathbf{L}^2 - \mathbf{F}^2$ , the effective Hamiltonian (36) takes the form

$$\hat{W}^{(0)} = \frac{p^2}{2m} + \hbar\omega_{\text{SOC}}(r) \left( \frac{\hat{\mathbf{J}}^2 - \mathbf{L}^2}{2\hbar^2} + \frac{1}{8} \right) + V_{\text{ex}}(\mathbf{r}), \quad (38)$$

where

$$\hat{\mathbf{J}} = \mathbf{L} + \hat{\mathbf{F}} \quad (39)$$

is the total angular momentum, and a uniform energy shift  $\hbar^2 \langle \mathcal{F}^2 \rangle / 2mr_0^2$  has been omitted in Eq. (38).

As one can see in Fig. 1, the SOC frequency  $\omega_{\text{SOC}}(r)$  decreases with the radius  $r$  and goes as  $r^{-2}$  for distances exceeding the SOC radius  $r_0$ :

$$\omega_{\text{SOC}}(r) \approx \frac{\hbar}{mr^2} \quad \text{for } r \gg r_0. \quad (40)$$

Such an asymptotic behavior of  $\omega_{\text{SOC}}(r)$  does not depend on the magnetic-field strength, and is determined exclusively by the ratio  $\hbar/m$ . The asymptotic Hamiltonian

$$\hat{W}^{(0)} = \frac{p^2}{2m} + \frac{1}{2mr^2} \left( \hat{\mathbf{J}}^2 - \mathbf{L}^2 + \frac{\hbar^2}{4} \right) + V_{\text{ex}}(\mathbf{r}) \quad (r \gg r_0) \quad (41)$$

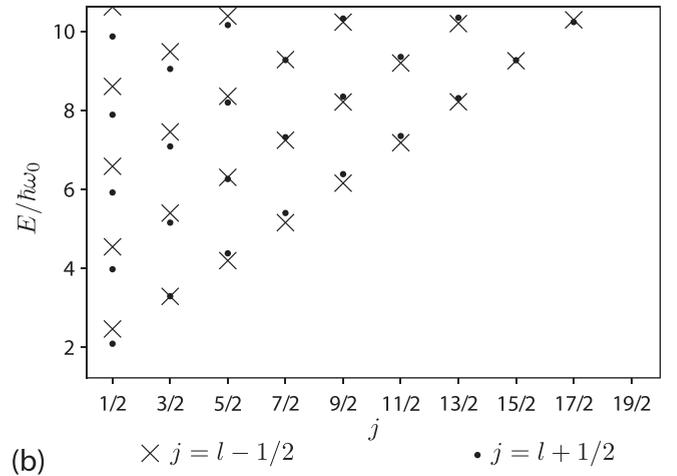
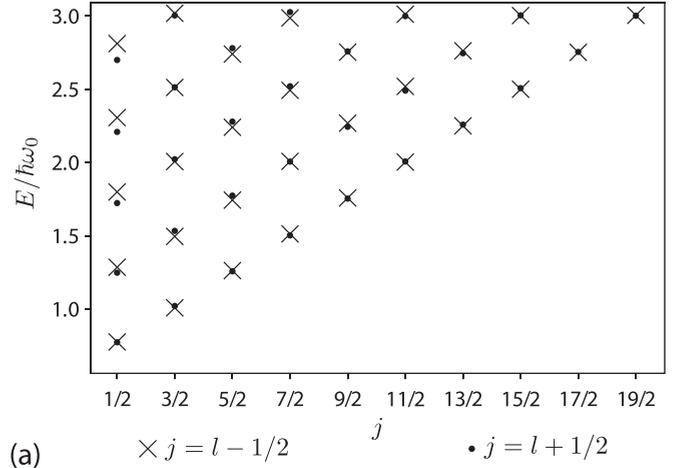


FIG. 2. Dependence of eigenenergies  $E_{n_r, j, l}$  on  $j$  for  $l = j \mp 1/2$  and up to five lowest radial quantum numbers  $n_r$ . The harmonic trapping potential (43) is added. Panel (a) corresponds to a softer trap with  $\eta = 0.5$ ; panel (b) corresponds to a tighter trap with  $\eta = 2$ . The spectrum is calculated for the sinusoidal (26) driving for which  $\langle \cos(2r\mathcal{F}/r_0) \rangle = \mathcal{J}_0(2r/r_0)$ .

contains a contribution  $\propto -\mathbf{L}^2$  which cancels the centrifugal term featured in the kinetic-energy operator

$$\frac{p^2}{2m} = -\frac{\hbar^2}{2m} \left[ \frac{1}{r^2} \frac{\partial}{\partial r} \left( r^2 \frac{\partial}{\partial r} \right) \right] + \frac{\mathbf{L}^2}{2mr^2}. \quad (42)$$

Thus, if the atomic eigenfunctions extend over distances exceeding the radius  $r_0$ , the corresponding eigenenergies are determined predominantly by the total angular momentum  $\hat{\mathbf{J}}^2$  and are nearly degenerate with respect to the orbital quantum number  $l = j \mp 1/2$  for fixed  $j$  and fixed radial quantum number  $n_r$ , showing a peculiar manifestation of the SOC. This is confirmed by numerical calculations presented in Sec. IV B 2 and displayed in Fig. 2. The long-range behavior of  $\omega_{\text{SOC}}(r)$  makes the SOC effects significant not only for the lower atomic states, but also for higher ones situated further away from the center. Note that for an electron in a Coulomb potential  $\propto -1/r$  the SOC strength is shorter ranged and goes as  $\propto r^{-3}$  [69], affecting mostly the lower electronic states situated closer to  $r = 0$ .

The SOC alone does not trap the atoms. Therefore, an external trapping potential  $V_{\text{ex}}(\mathbf{r})$  is needed to have bound states, like in the case of the light induced geometric potentials [80,89]. The external potential can be chosen freely. For example, one can take  $V_{\text{ex}}(\mathbf{r})$  to be a spherical harmonic trapping potential

$$V_{\text{ex}}(\mathbf{r}) = \frac{m}{8}\eta^2\omega_0^2r^2, \quad (43)$$

where  $\eta$  defines the trapping frequency  $\omega_{\text{ex}} = \eta\omega_0/2$  and  $\omega_0 = \omega_{\text{SOC}}(0)$  is the SOC frequency at zero distance. For the sinusoidal driving (26) one has

$$\omega_0 = \frac{\hbar}{mr_0^2}. \quad (44)$$

## 2. Angular states

In what follows we shall consider a spherically symmetric external potential  $V_{\text{ex}}(\mathbf{r}) = V_{\text{ex}}(r)$ . The Hamiltonian  $\hat{W}^{(0)}$  given by Eq. (38) contains the commuting operators  $\hat{\mathbf{J}}^2$  and  $\mathbf{L}^2$  characterized by the eigenvalues  $\hbar^2j(j+1)$  and  $\hbar^2l(l+1)$ . The eigenstates  $|j, l, f, m_j\rangle$  of  $\hat{W}^{(0)}$  are thus described by the quantum numbers  $j, l, f$ , and  $m_j$ , with  $l - f \leq j \leq l + f$  and  $f = 1/2$ . The eigenstates  $|j, l, f, m_j\rangle = |j, l, f, m_j(\theta, \phi)\rangle$  are degenerate with respect to projection of the total angular momentum  $-j \leq m_j \leq j$ . They can be cast in terms of the angular momentum states  $Y_{l,m}(\theta, \phi)$  (the spherical harmonics) and the spin states  $|f, m_f\rangle$  with  $f = 1/2$ :

$$|j, l, f, m_j(\theta, \phi)\rangle = \sum_{m_l=-l}^l \sum_{m_f=\pm\frac{1}{2}} Y_{l,m_l}(\theta, \phi) |f, m_f\rangle \times \langle l, m_l, f, m_f | j, m_j \rangle, \quad (45)$$

where  $\theta$  and  $\phi$  are the polar and azimuthal angles,  $\langle l, m_l, f, m_f | j, m_j \rangle$  is the Clebsch-Gordan coefficient, and the summation is over the projections of the spin and the orbital angular momentum with  $m_l + m_f = m_j$ .

## 3. Radial eigenequation

The full eigenstate of the effective Hamiltonian (38) contains also the radial part

$$|\psi_{n_r, j, l, f}(r, \theta, \phi)\rangle = |j, l, f, m_j(\theta, \phi)\rangle \psi_{n_r, j, l, f}(r), \quad (46)$$

where  $n_r$  is a radial quantum number. Substituting

$$\psi_{n_r, j, l, f, m_j}(r) \equiv \frac{\phi_{n_r, j, l, f, m_j}(r)}{r}, \quad (47)$$

one arrives at the eigenvalue equation for the scaled radial function  $\phi_{n_r, j, l, f, m_j}(r)$

$$\left[ -\frac{\hbar^2}{2m} \frac{\partial^2}{\partial r^2} + V_{j,l}(r) \right] \phi_{n_r, j, l, f, m_j}(r) = E_{n_r, j, l} \phi_{n_r, j, l, f, m_j}(r), \quad (48)$$

subject to the condition  $\phi_{n_r, j, l, f, m_j}(0) = 0$ , where

$$V_{j,l}(r) = \frac{1}{2} \left[ j(j+1) - l(l+1) + \frac{1}{4} \right] \hbar\omega_{\text{SOC}}(r) + \frac{\hbar^2 l(l+1)}{2mr^2} + V_{\text{ex}}(r) \quad (49)$$

is the radial potential and  $E_{n_r, j, l}$  is an eigenenergy.

## 4. Analysis of eigenenergies

Figure 2 displays the dependence of the eigenenergies  $E_{n_r, j, l}$  on  $j$  for  $l = j \mp 1/2$  and up to five lowest radial quantum numbers  $n_r$ . The calculations are carried out for the sinusoidal driving, Eq. (26), and the harmonic trapping potential given by Eq. (43) with  $\eta = 0.5$  and  $\eta = 2$ . For a softer trap ( $\eta = 0.5$ ) and  $j \geq 3/2$ , there is an almost perfect degeneracy of the eigenenergies  $E_{n_r, j, l}$  with the same  $j$  and  $n_r$  but different  $l = j \mp 1/2$ . For such a softer trap the atomic wave functions extend to distances  $r \gg r_0$  in which  $\omega_{\text{SOC}}(r) \approx \hbar/mr^2$ , corresponding to the strong driving regime ( $g_F B/2 \gg \omega$ ). Consequently the  $l$ -dependent part of the SOC term cancels the centrifugal term in Eq. (49), and the eigenstates depend weakly on  $l$ . For a tighter trap ( $\eta = 2$ ) the atom is localized closer to the center leading to a larger difference in the eigenenergies  $E_{n_r, j, l}$  with different  $l = j \mp 1/2$  but the same  $j$  and  $n_r$ .

It is noteworthy that an infinite set of degenerate eigenstates (3D Landau levels [90]) is formed if a particle is subjected to the SOC term  $\pm\omega_C \mathbf{L} \cdot \hat{\boldsymbol{\sigma}}$  with a constant frequency  $\omega_C$ , and a 3D isotropic harmonic trap is added with the frequency  $\omega_C$  [90–92]. The eigenstates with  $j = l \mp 1/2$  then have eigenenergies  $\hbar\omega_C(2n_r + 1 \mp 1/2)$ , which depend only on the principal quantum number  $n_r$  and thus are degenerate with respect to the  $l$  and  $j$ . In the present study the situation is different. The SOC frequency  $\omega_{\text{SOC}}(r)$  decreases with the radius and has a special asymptotic behavior at large distances,  $\omega_{\text{SOC}}(r) \approx \hbar/mr^2$ , leading to pairs of close energy levels with  $l = j \pm 1/2$  for fixed  $j$ , as discussed above.

For the sinusoidal driving, Eq. (26), the difference in the radial potentials with  $l = j \pm 1/2$  for fixed  $j$ ,

$$\begin{aligned} \Delta V_j(r) &= V_{j, l=j+1/2}(r) - V_{j, l=j-1/2}(r) \\ &= \frac{\hbar(2j+1)\mathcal{J}_0(2r/r_0)}{mr^2}, \end{aligned} \quad (50)$$

is determined by the Bessel function  $\mathcal{J}_0(2r/r_0)$ , which is positive for distances  $2r/r_0$  smaller than 2.4. Consequently, the ground state with  $j = 1/2$  and  $l = 0$  has a slightly lower energy than the one with  $j = 1/2$  and  $l = 1$ , as one can see in Fig. 2. The situation can be reversed by adding an extra antitrapping potential for small  $r$ . Such a potential pushes the atomic probability distribution to a region of larger distances,  $2r/r_0 > 2.4$ , where the Bessel function  $\mathcal{J}_0(2r/r_0)$  becomes negative and reaches the maximum negative value of  $-0.36$  at  $2r/r_0 = 3.83$ . Figure 3 shows the difference in the ground-state energies  $\Delta E = E_{j=\frac{1}{2}, l=1} - E_{j=\frac{1}{2}, l=0}$  for the external trapping potential  $V_{\text{ex}}(r)$  composed of a spherically symmetric harmonic potential with  $\eta = 0.5$  (upper plot) or  $\eta = 2$  (lower plot), and an additional hard-core potential of a radius  $r = r_*$  preventing the atom from being at distances  $r \leq r_*$ . For  $\eta = 0.5$  ( $\eta = 2$ ) the energy difference  $\Delta E$  becomes negative at  $r_*/r_0 = 0.14$  ( $r_*/r_0 = 0.55$ ) and reaches the maximum negative value of  $\Delta E = -0.019\hbar\omega_0$  ( $\Delta E = -0.084\hbar\omega_0$ ) at  $r_*/r_0 = 0.55$  ( $r_*/r_0 = 0.97$ ).

In this way, the ground state of the system can be the state with  $j = 1/2$  and the orbital quantum number equal to  $l = 1$  rather than  $l = 0$ . For conventional spherically symmetric

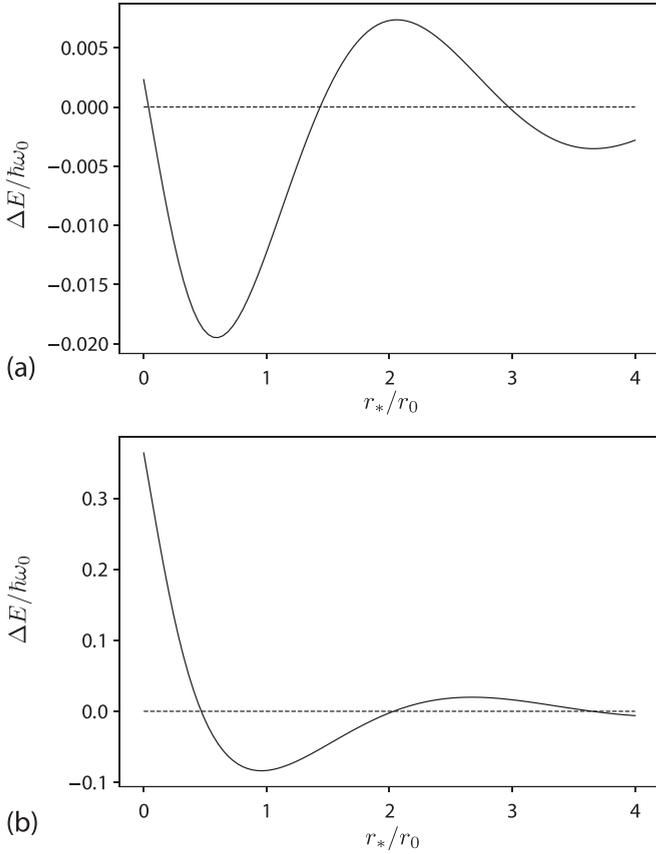


FIG. 3. Energy difference  $\Delta E = E_{j=\frac{1}{2},l=1} - E_{j=\frac{1}{2},l=0}$  vs the radius  $r_*$  of the additional hard core potential for an atom in a harmonic trapping potential (43) with  $\eta = 0.5$  [panel (a)] and  $\eta = 2$  [panel (b)]. The calculations are done for the sinusoidal driving (26) for which  $\langle \cos(2r\mathcal{F}/r_0) \rangle = \mathcal{J}_0(2r/r_0)$ .

systems, such as the hydrogenlike atoms, the ground state is always characterized by  $l = 0$  and thus is not affected by the SOC. The formation of the ground state with  $l = 1$  is now facilitated by the time-periodic driving which induces a longer-ranged SOC  $\propto 1/r^2$  and allows one to reverse the sign of the potential difference  $\Delta V_j(r)$  in Eq. (50) due to the change of sign in the Bessel function  $\mathcal{J}_0(2r/r_0)$ . Note that the periodic driving can be also used to reverse the sign of the matrix elements of tunneling of atoms in optical lattices [17,93–95]. By shaking the lattice sufficiently strongly, the Bessel function renormalizes the tunneling matrix elements, causing them to change the sign [17,94,95].

## V. ADIABATIC CONDITIONS AND IMPLEMENTATION

### A. Adiabatic condition

The general adiabatic condition (11) was discussed in Sec. II C. Now we will consider in more detail the adiabatic condition for an atom in the magnetic field. Using Eqs. (7), (11), (B1), and (B5), one arrives at the adiabatic condition for the atom in a centrally symmetric magnetic field

$$\omega \gg \omega_0, \quad \omega \gg \omega_{\text{kin}}, \quad (51)$$

where  $\omega_0 = \hbar/mr_0^2$  is the SOC frequency and  $\hbar\omega_{\text{kin}}$  is the kinetic energy of the atomic motion. The first requirement in Eq. (51) is due to the  $[\hat{\mathbf{A}}(\mathbf{r}, \omega t)]^2$  term in the transformed Hamiltonian  $\hat{W}$ , Eq. (7), with the second one coming from the mixed terms  $\hat{\mathbf{A}}(\mathbf{r}) \cdot \mathbf{p}$  and  $\mathbf{p} \cdot \hat{\mathbf{A}}(\mathbf{r})$ . A condition similar to Eq. (51) can be obtained also for a more general cylindrically symmetric magnetic field given by Eq. (31).

The adiabatic condition (51) does not rely on the smallness of the frequency of the magnetic interaction  $g_F B(r)/2 = \omega r/r_0$  compared to the driving frequency  $\omega$ , so we do not restrict ourselves to distances smaller than the SOC radius  $r_0$ . We only require the SOC frequency  $\omega_0$  and the atomic kinetic frequency  $\omega_{\text{kin}}$  to be much smaller than the driving frequency  $\omega$ . Hence the present approach allows one to study the SOC at distances exceeding the ones accessible via the perturbative treatment in the original representation. Specifically, the SOC frequency  $\omega_0 = \hbar/mr_0^2$  can now be considerably larger than the one accessible by means of the original perturbation approach [74] applicable if the radius of the atomic cloud is much smaller than the SOC radius  $r_0$ .

The frequency of magnetic interaction  $g_F B(r)/2 = \omega r/r_0$  reaches the driving frequency at  $r = r_0$ . Consequently, one can keep  $r_0$  fixed by simultaneously increasing both the magnetic-field strength and the driving frequency until the latter  $\omega$  becomes sufficiently large compared to the SOC frequency  $\omega_0$  to fulfill the adiabatic condition (51).

In many-body systems there can be additional losses due to atom-atom interactions. The two-body losses have been studied recently for a periodic driving of the form  $\hbar k_0 \sigma_z z \cos(\omega t)$  [96]. For the driving energy  $\hbar\omega$  exceeding both the kinetic energy and the SOC energy  $\hbar^2 k_0^2/2m$ , the two-body heating rate was shown to increase as  $\sqrt{\hbar\omega}$  due to an increase of the final density of states and the energy of the quantum absorbed  $\hbar\omega$  [96]. Yet the probability of such absorption events is proportional to  $1/\sqrt{\hbar\omega}$  and thus goes to zero in the limit of an infinitely large driving frequency. Thus the many-body heating can be minimized by increasing the driving frequency and removing from the trap a handful of very fast atoms which absorb the driving quantum  $\hbar\omega$ . This is essentially the idea of evaporative cooling [97]. If  $\hbar\omega$  exceeds the trap depth, then those few atoms absorbing the quantum of  $\hbar\omega$  are automatically ejected from the trap which becomes shallow at large energies or large distances. Such a trap can be produced optically by focusing a number of laser beams within the atomic cloud. The many-body effects will be explored in more detail in a separate study.

### B. Implementation

In analyzing the 3D SOC induced by the oscillating magnetic field  $\mathbf{B}(\mathbf{r})f(\omega t)$  we used cylindrically or spherically symmetric magnetic fields with the amplitude  $\mathbf{B}(\mathbf{r})$  given by Eqs. (31) or (33). Such a magnetic field generally has a nonzero divergence and thus does not obey the Maxwell equation. Yet one can produce interaction between the atom and the magnetic field characterized by an effective magnetic field with a nonzero divergence. In particular, one can generate the spherically symmetric monopole field given by Eq. (33). As explained in detail in Appendix A, this can be

done by taking the actual (real) magnetic field  $\mathbf{B}_{\text{real}}$  entering the original interaction operator  $\hat{V}_{\text{real}}(\mathbf{r}, t) = g_F \hat{\mathbf{F}} \cdot \mathbf{B}_{\text{real}}(\mathbf{r}, t)$  to contain a time-independent bias magnetic field  $B_0 \mathbf{e}_z$  and an extra spatially inhomogeneous magnetic field proportional to the time-periodic function  $f(\omega t)$ :

$$\mathbf{B}_{\text{real}}(\mathbf{r}, t) = B_0 \mathbf{e}_z + [\mathbf{B}_1(\mathbf{r}) + \mathbf{B}_2(\mathbf{r}) \cos(\omega_B t)] f(\omega t), \quad (52)$$

where the frequency  $\omega_B$  is considered to be in an exact resonance with magnetic level splitting induced by the bias field:  $\omega_B = g_F B_0$ . Furthermore,  $\omega_B$  is taken to be much larger than the frequency of the periodic driving:  $\omega_B \gg \omega$ . Transforming the spin to the frame rotating at the frequency  $\omega_B$  via the unitary transformation  $U$  given by Eq. (A4), one can then apply the rotating wave approximation (RWA) and neglect the fast oscillating terms  $\propto \exp(\pm i\omega_B t)$  and  $\propto \exp(\pm 2i\omega_B t)$  in the transformed interaction operator. Consequently one arrives at the interaction of the spin with the time-periodic effective magnetic field  $\mathbf{B}(\mathbf{r})f(\omega t)$  characterized by the amplitude (see Appendix A for more details)

$$\mathbf{B}(\mathbf{r}) = \frac{1}{2} B_{2x}(\mathbf{r}) \mathbf{e}_x + \frac{1}{2} B_{2y}(\mathbf{r}) \mathbf{e}_y + B_{1z}(\mathbf{r}) \mathbf{e}_z. \quad (53)$$

In this way, one makes use of two unitary transformations. The transformation  $U$  given by Eq. (A4) eliminates the fast spin precession around the bias magnetic field at the frequency  $\omega_B$ . This provides an effective coupling of the atom with the effective magnetic field given by Eq. (53). Subsequently one applies another unitary transformation  $R$  given by Eq. (4) which eliminates the interaction operator  $g_F \hat{\mathbf{F}} \cdot \mathbf{B}(\mathbf{r})f(\omega t)$ . The position dependence of the unitary operator  $R = R(\mathbf{r}, t)$  yields then the SOC due to the spin-dependent momentum shift in the kinetic-energy term of Eq. (7).

Using such an approach one can create various effective magnetic fields  $\mathbf{B}(\mathbf{r})$  which do not necessarily obey the Maxwell equations. In particular, by taking  $\mathbf{B}_1(\mathbf{r})$  and  $\mathbf{B}_2(\mathbf{r})$  to be the quadrupole magnetic fields, one generates the cylindrically symmetric field or the spherically symmetric monopole field given by Eqs. (31) and (33), respectively; see Appendix A and Ref. [70].

The method works if the frequency of the magnetic level splitting  $\omega_B$  is much larger than the driving frequency  $\omega$ . In the experiment [98] with  $^{87}\text{Rb}$  atoms, the magnetic splitting frequency  $\omega_B = 2\pi \times 4.81$  MHz far exceeds the recoil frequency  $\omega_{\text{rec}}$  which equals  $2\pi \times 3.77$  kHz for the 780 nm  $5^2S_{1/2} \rightarrow 5^2P_{3/2}$  optical transition. Note that the bias magnetic field induces also the quadratic Zeeman shift (QZS) with a frequency equal approximately to  $6\omega_{\text{rec}}$  in the experiment [98]. The unwanted QZS can be reduced by decreasing the bias magnetic field. For example, by reducing  $\omega_B$  to  $2\pi \times 50$  kHz, the QZS decreases to  $0.001\omega_{\text{rec}}$ , which is in the range of the few Hz and thus can be completely neglected.

On the other hand, QZS can be used to produce an effective quasi-spin-1/2 system for atoms characterized by larger spins [41,99]. For this the oscillating magnetic field should be in resonance with a selected pair of magnetic levels, and the QZS makes the coupling with other spin states out of resonance. For example, the magnetic field could resonantly couple the  $m_F = -1$  and  $m_F = 0$  states of the  $F = 1$  manifold of  $^{87}\text{Rb}$  or  $^{23}\text{Na}$  atoms [41] representing the quasi-spin-up and -down states, leaving the detuned  $m_F = 1$  state uncoupled, similar to experiments on the Raman-induced SOC [98]. To create

the quasi-spin-1/2 system, the driving frequency  $\omega$  should be smaller than the frequency of the quadratic Zeeman shift  $\omega_q$ . As mentioned above,  $\omega_q$  equals a few recoil frequencies in the experiment [98]. The QZS can be further increased by increasing the bias magnetic field to reach the condition  $\omega_q \gg \omega_{\text{rec}}$ , so the driving frequency  $\omega$  can be of the order of the recoil frequency  $\omega_{\text{rec}}$  or a little above it. Therefore, the SOC frequency  $\omega_0 = \hbar/mr_0^2$  should then be smaller than the recoil frequency to fulfill the adiabatic requirements (51). For example, for  $^{87}\text{Rb}$  atoms,  $\omega_0$  could be of the order of a few tens to a few hundreds of Hz, which is comparable to typical trapping frequencies. This provides the SOC radius  $r_0$  of the order of a few optical wavelengths. In this way, by taking a width of an optical trap to be of the order of ten optical wavelengths, the SOC radius  $r_0$  would be within the trapped atomic cloud, and the atoms would experience a substantial SOC in the cloud.

## VI. CONCLUDING REMARKS

We have considered a method of creating the non-Abelian geometric potential and thus the 3D SOC for the center-of-mass motion of the particle subjected to periodic driving. The periodic perturbation is a product of a position-dependent Hermitian operator  $\hat{V}(\mathbf{r})$  and a fast oscillating periodic function  $f(\omega t)$  with a zero average. To have a significant SOC, we have analyzed a situation where the matrix elements of the periodic operator  $\hat{V}(\mathbf{r})f(\omega t)$  are not necessarily small compared to the driving energy  $\hbar\omega$ , so that one cannot apply the high frequency expansion of the effective Floquet Hamiltonian [56,71–73] in the original representation. To bypass the problem, we have applied a unitary transformation which eliminates the original periodic perturbation and yields an oscillating vector potential term. The resulting periodic perturbation is no longer proportional to the driving frequency  $\omega$ , so the perturbation treatment is applicable to much stronger driving (and thus over a larger range of distances) than in the original representation. We have considered a situation where  $\hat{V}(\mathbf{r})$  depends on internal (spin or quasispin) degrees of freedom of the particle, and thus the operator  $\hat{V}(\mathbf{r})$  does not necessarily commute with itself at different positions:  $[\hat{V}(\mathbf{r}), \hat{V}(\mathbf{r}')] \neq 0$ . Consequently, the adiabatic evolution of the system within a Floquet band is accompanied by a non-Abelian (noncommuting) geometric vector potential providing the 3D SOC.

The periodic driving plays a vital role in our analysis. Without the periodic driving the interaction is given by a time-independent operator  $V(\mathbf{r})$ , like in Ref. [70]. In that case the spin adiabatically follows the magnetic field and is fully polarized along local field direction [70]. In the present situation, the interaction operator  $V(\mathbf{r})f(\omega t)$  contains also the periodic function  $f(\omega t)$  with the zero average. Therefore, the spin no longer adiabatically follows the magnetic field and thus is no longer polarized along the local magnetic field. Specifically, the operator  $V(\mathbf{r})f(\omega t)$  provides fully degenerate Floquet bands [57]. The position dependence of these Floquet eigenstates yields the spin-dependent momentum shift and thus the spin-orbit coupling. Therefore, the current situation is very different from the case of the time-independent

interaction  $V(\mathbf{r})$ , where the eigenenergies of  $V(\mathbf{r})$  are nondegenerate and thus the spin is polarized [70].

The general formalism has been illustrated by analyzing motion of a spinful atom in a magnetic field oscillating in time, subsequently concentrating on a spin-1/2 atom in a cylindrically symmetric magnetic field. This yields the SOC involving coupling between the spin and the orbital motion described by all three components of the OAM operator  $\mathbf{L}$ . In particular, the time-oscillating monopole-type magnetic field  $\mathbf{B} \propto \mathbf{r}$  generates the 3D SOC of the  $\mathbf{L} \cdot \mathbf{F}$  form. The strength of this SOC goes as  $1/r^2$  for larger distances, rather than  $1/r^3$ , as for electrons in the Coulomb potential. Such a long-ranged SOC significantly affects not only the lower states of the trapped atom, but also the higher ones. In particular, the states with  $l = j \pm 1/2$  are nearly degenerate with fix  $j$  and  $n_r$  for an atom characterized by the (quasi)spin 1/2. In the presence of a harmonic external trapping potential, the ground state with  $j = 1/2$  and  $l = 0$  has a slightly lower energy than the one with  $j = 1/2$  and  $l = 1$ . The situation can be reversed by adding an extra antitrapping potential for small  $r$ , which causes the ground state of the system to be characterized by the orbital quantum number  $l = 1$ . The  $l = 1$  ground state is affected by the SOC, which can lead to interesting many-body phases to be explored elsewhere. If the atom possesses higher spin, more complicated SOC terms can be generated. In this situation, the spin-dependent scalar potential featured in Eq. (36) for the effective Hamiltonian  $\hat{W}^{(0)}$  can lead to an additional coupling between orbit and spin or spin tensor. We plan to address this topic in a future study.

Previously coupling between the spin and the linear momentum  $\mathbf{p}$  was considered for ultracold atoms using time-periodic sequences of magnetic pulses [35,36,40,41]. To generate the 2D or 3D coupling between the spin and linear momentum  $\mathbf{p}$ , the periodic driving involves rapid changes of the magnetic-field direction [35,36]. This would be rather complicated to implement experimentally. It is much more straightforward to generate a sizable coupling between the spin and the OAM using the method considered here. For this one applies a simpler magnetic field with properly designed spatial and temporal profiles rather than the alternating magnetic pulses. Therefore, the current scheme is more realistic and can be implemented using experimental techniques currently available.

The 2D and 3D SOC can be also generated optically by using a degeneracy of eigenstates of the atom-light coupling operator [59–68], which involves a considerable amount of effort [64,66–68]. The present approach does not require such a degeneracy. Instead, the periodic driving yields degenerate Floquet states for the time-periodic interaction operator  $\hat{V}(\mathbf{r})f(\omega t)$  in a straightforward way [57]. The spatial and temporal dependence of the operator  $\hat{V}(\mathbf{r})f(\omega t)$  provides the oscillating vector potential and hence the SOC. The present SOC has also some similarities to the coupling between the spin and OAM induced by Raman laser beams carrying optical vortices [76–87]. Yet, unlike the Raman case, now the coupling between the spin and OAM is described by all three OAM projections  $L_x$ ,  $L_y$ , and  $L_z$ , so the SOC is truly three dimensional. Furthermore, our current scheme does not involve laser fields, and hence does not suffer from Raman-induced heating.

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## APPENDIX A: GENERATION OF EFFECTIVE MAGNETIC FIELD VIOLATING THE MAXWELL EQUATIONS

In this Appendix we will explain a way of obtaining the coupling of an atom with an effective magnetic field which does not necessarily have a zero divergence and/or a zero curl [35,41,70]. In particular, the effective magnetic field can describe a monopole-type field [70].

### 1. Original problem

In the original problem, the atom is coupled with an actual (real) magnetic field  $\mathbf{B}_{\text{real}}(\mathbf{r}, t)$  obeying the Maxwell equations. The corresponding interaction operator reads

$$\hat{V}_{\text{real}}(\mathbf{r}, t) = g_F \hat{\mathbf{F}} \cdot \mathbf{B}_{\text{real}}(\mathbf{r}, t). \quad (\text{A1})$$

Let us consider a magnetic field

$$\mathbf{B}_{\text{real}}(\mathbf{r}, t) = B_0 \mathbf{e}_z + [\mathbf{B}_1(\mathbf{r}) + \mathbf{B}_2(\mathbf{r}) \cos(\omega_B t)] f(\omega t), \quad (\text{A2})$$

which is composed of a constant bias magnetic field  $B_0 \mathbf{e}_z$  pointing along a unit Cartesian vector  $\mathbf{e}_z$ , and a time-dependent term proportional to the periodic function  $f(\omega t) = f(\omega t + 2\pi)$  oscillating in time with the frequency  $\omega$ . The latter term contains also a contribution  $\propto \cos(\omega_B t)$  oscillating with a frequency  $\omega_B$  much larger than the driving frequency  $\omega$ . The frequency  $\omega_B$  is taken to be in an exact resonance with the frequency of the splitting between the spin states induced by the bias magnetic field, so that

$$\omega_B = g_F B_0, \quad \text{with } \omega_B \gg \omega. \quad (\text{A3})$$

Replacing  $\cos(\omega_B t)$  by  $\cos(\omega_B t + \phi)$  one can also include the phase shift  $\phi$  of the fast oscillations in Eq. (A2). This can be done within the present framework by changing the origin of time  $t \rightarrow t + \phi/\omega_B$  and redefining the periodic function:  $f(\omega t) \rightarrow f(\omega t - \phi\omega/\omega_B)$ .

### 2. Transformed representation

The bias magnetic field can be eliminated via a unitary transformation which rotates the spin at the frequency  $\omega_B$  around the  $z$  axis

$$\hat{U} = \exp(-i\omega_B t \hat{F}_z / \hbar). \quad (\text{A4})$$

The transformed Hamiltonian

$$\hat{V}_{\text{transf}}(\mathbf{r}, t) = \hat{U}^\dagger \hat{V}_{\text{real}}(\mathbf{r}, t) \hat{U} - i\hbar \hat{U}^\dagger \partial_t \hat{U} \quad (\text{A5})$$

no longer contains the bias term  $g_F B_0 \hat{F}_z$  which is canceled by the term  $-i\hbar \hat{U}^\dagger \partial_t \hat{U}$ , giving

$$\hat{V}_{\text{transf}}(\mathbf{r}, t) = g_F \hat{\mathbf{F}} \cdot [\mathbf{B}_1(\mathbf{r}) + \mathbf{B}_2(\mathbf{r}) \cos(\omega_B t)] f(\omega t). \quad (\text{A6})$$

Here  $\hat{\mathbf{F}} = \hat{U}^\dagger \hat{\mathbf{F}} \hat{U}$  is the transformed spin operator. Its  $z$  component  $\hat{F}_z = \hat{F}_z$  is not affected by the transformation. On the other hand, the  $x$  and  $y$  components of  $\hat{\mathbf{F}}$  rotate at the frequency  $\omega_B$ :

$$\hat{F}_x = \hat{F}_x \cos(\omega_B t) - \hat{F}_y \sin(\omega_B t), \quad (\text{A7})$$

$$\hat{F}_y = \hat{F}_x \sin(\omega_B t) + \hat{F}_y \cos(\omega_B t). \quad (\text{A8})$$

### 3. Effective interaction operator and effective magnetic field

Since  $\omega_B \gg \omega$ , one can apply the rotating wave approximation (RWA) replacing the fast oscillating term  $\hat{\mathbf{F}} \cdot \mathbf{B}_1(\mathbf{r}) f(\omega t)$  by its temporal average  $\hat{F}_z B_{1z}(\mathbf{r}) f(\omega t)$  in Eq. (A6). In a similar manner, the fast oscillating term  $\hat{\mathbf{F}} \cdot \mathbf{B}_2(\mathbf{r}) \cos(\omega_B t)$  averages to  $\frac{1}{2} \hat{F}_x B_{2x}(\mathbf{r}) + \frac{1}{2} \hat{F}_y B_{2y}(\mathbf{r})$  in Eq. (A6). Therefore, by averaging  $\hat{V}_{\text{transf}}(\mathbf{r}, t)$  over a small period  $T_B = 2\pi/\omega_B$ , one arrives at an effective interaction operator

$$\langle \hat{V}_{\text{transf}}(\mathbf{r}, t) \rangle = g_F \hat{\mathbf{F}} \cdot \mathbf{B}(\mathbf{r}) f(\omega t), \quad (\text{A9})$$

where

$$\mathbf{B}(\mathbf{r}) = \frac{1}{2} B_{2x}(\mathbf{r}) \mathbf{e}_x + \frac{1}{2} B_{2y}(\mathbf{r}) \mathbf{e}_y + B_{1z}(\mathbf{r}) \mathbf{e}_z \quad (\text{A10})$$

is a time-independent amplitude of the oscillating effective magnetic field.

Although the original fields  $\mathbf{B}_1(\mathbf{r})$  and  $\mathbf{B}_2(\mathbf{r})$  entering Eq. (A2) are divergence free, the effective field  $\mathbf{B}(\mathbf{r})$  featured in Eq. (A9) does not necessarily obey this requirement. For example, by taking  $\mathbf{B}_1(\mathbf{r}) = 0$  and  $\mathbf{B}_2(\mathbf{r}) \propto x\mathbf{e}_x - z\mathbf{e}_z$ , one arrives at an effective magnetic field pointing along the  $x$  axis  $\mathbf{B}(\mathbf{r}) \propto x\mathbf{e}_x$  with  $\nabla \cdot \mathbf{B}(\mathbf{r}) \neq 0$  [35].

### 4. Cylindrical and monopole fields

Suppose now that both  $\mathbf{B}_1(\mathbf{r})$  and  $\mathbf{B}_2(\mathbf{r})$  entering the original magnetic field (A2) are the quadrupole fields:

$$\mathbf{B}_1(\mathbf{r}) = -\frac{1}{2} \alpha_z (x\mathbf{e}_x + y\mathbf{e}_y - 2z\mathbf{e}_z), \quad (\text{A11})$$

$$\mathbf{B}_2(\mathbf{r}) = 2\alpha_\perp (x\mathbf{e}_x + y\mathbf{e}_y - 2z\mathbf{e}_z), \quad (\text{A12})$$

where  $\alpha_z$  and  $\alpha_\perp$  characterize the strengths of the constituting magnetic fields. In that case the effective magnetic field (A2) generally contains all three Cartesian components and is given by

$$\mathbf{B}(\mathbf{r}) = \alpha_\perp (x\mathbf{e}_x + y\mathbf{e}_y) + \alpha_z z \mathbf{e}_z. \quad (\text{A13})$$

The field  $\mathbf{B}(\mathbf{r})$  has a nonzero divergence  $\nabla \cdot \mathbf{B}(\mathbf{r}) \neq 0$  as long as  $\alpha_z \neq -2\alpha_\perp$ . In particular, for  $\alpha_z = \alpha_\perp = \alpha$  one arrives at the monopole-type [70] effective magnetic field for which  $\mathbf{B}(\mathbf{r}) = \alpha \mathbf{r}$ .

### APPENDIX B: CALCULATION OF $\langle [\hat{\mathbf{A}}(\mathbf{r}, \omega t)]^2 \rangle$ TERM

Let us analyze the  $\langle [\hat{\mathbf{A}}(\mathbf{r}, \omega t)]^2 \rangle$  term featured in Eq. (14) for the scalar potential  $\hat{V}_{\text{total}}(\mathbf{r})$ . For this let us rewrite Eq. (27) in vector notations:

$$\hat{\mathbf{A}}(\mathbf{r}, \omega t) = a\mathcal{F} \mathbf{d}_1 - \sin(a\mathcal{F}) \mathbf{d}_2 - [\cos(a\mathcal{F}) - 1] \mathbf{d}_3, \quad (\text{B1})$$

where  $\mathbf{d}_1$ ,  $\mathbf{d}_2$ , and  $\mathbf{d}_3$  are vectors with the Cartesian components:

$$d_{1u} = \frac{(\mathbf{B} \cdot \mathbf{F})(\mathbf{B} \cdot \partial \mathbf{B} / \partial u)}{B^3}, \quad (\text{B2})$$

$$d_{2u} = \frac{[\mathbf{B} \times (\mathbf{B} \times \mathbf{F})] \cdot \partial \mathbf{B} / \partial u}{B^3}, \quad (\text{B3})$$

$$d_{3u} = \frac{(\mathbf{B} \times \mathbf{F}) \cdot \partial \mathbf{B} / \partial u}{B^2}. \quad (\text{B4})$$

Since  $\mathcal{F}(\omega t)$  is an odd function of time  $\mathcal{F}(\omega t) = -\mathcal{F}(-\omega t)$ , the terms containing  $\mathbf{d}_1 \cdot \mathbf{d}_3$  and  $\mathbf{d}_2 \cdot \mathbf{d}_3$  average to zero in  $\langle [\hat{\mathbf{A}}(\mathbf{r}, \omega t)]^2 \rangle$ . Furthermore, for the spherically symmetric monopole magnetic field with  $\mathbf{B} = \alpha \mathbf{r}$ , one has

$$\mathbf{d}_1 = \frac{(\mathbf{r} \cdot \mathbf{F}) \mathbf{r}}{r^3}, \quad \mathbf{d}_2 = \frac{\mathbf{r} \times (\mathbf{r} \times \mathbf{F})}{r^3}, \quad \mathbf{d}_3 = \frac{(\mathbf{r} \times \mathbf{F})}{r^2}, \quad (\text{B5})$$

giving  $\mathbf{d}_1 \cdot \mathbf{d}_2 = 0$ , as well as

$$d_1^2 = \frac{(\mathbf{r} \cdot \mathbf{F})^2}{r^4}, \quad d_2^2 = d_3^2 = \frac{F^2 r^2 - (\mathbf{r} \cdot \mathbf{F})^2}{r^4}. \quad (\text{B6})$$

Consequently one arrives at

$$\langle [\hat{\mathbf{A}}(\mathbf{r}, \omega t)]^2 \rangle = a^2 \langle \mathcal{F}^2 \rangle d_1^2 + 2d_2^2 [1 - \langle \cos(a\mathcal{F}) \rangle]. \quad (\text{B7})$$

Equations (B6) and (B7) provide the following explicit result for the monopole field:

$$\begin{aligned} \langle [\hat{\mathbf{A}}(\mathbf{r}, \omega t)]^2 \rangle &= \frac{4}{r_0^2} \langle \mathcal{F}^2 \rangle \frac{(\mathbf{r} \cdot \mathbf{F})^2}{r^2} \\ &+ \frac{2 - 2\langle \cos(2r\mathcal{F}/r_0) \rangle}{r^2} \frac{F^2 r^2 - (\mathbf{r} \cdot \mathbf{F})^2}{r^2}. \end{aligned} \quad (\text{B8})$$

The relation of the form (B7) holds also for the spin-1/2 atom ( $\mathbf{F} = \frac{\hbar}{2} \boldsymbol{\sigma}$ ) in a cylindrically symmetric magnetic field (31) for which  $d_2^2 = d_3^2$  and  $\mathbf{d}_1 \cdot \mathbf{d}_2 + \mathbf{d}_2 \cdot \mathbf{d}_1 = 0$ . In such a situation

$$d_1^2 = \frac{\hbar^2 (\alpha_\perp^4 \rho^2 + \alpha_z^4 z^2)}{4 (\alpha_z^2 z^2 + \alpha_\perp^2 \rho^2)^2} \quad (\text{B9})$$

and

$$d_2^2 = \frac{\hbar^2 \alpha_\perp^2 (2\alpha_z^2 z^2 + (\alpha_z^2 + \alpha_\perp^2) \rho^2)}{4 (\alpha_z^2 z^2 + \alpha_\perp^2 \rho^2)^2}. \quad (\text{B10})$$

This is consistent with the spherically symmetric result (B6) for the spin-1/2 atom.

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